

## Photodesorption of X-ray Irradiated H<sub>2</sub>O:CO:NH<sub>3</sub> Ice Mixtures

N.-E. Sie (謝妮恩)<sup>1\*</sup>, A. Jiménez-Escobar<sup>2</sup>, A. Ciaravella<sup>2</sup>, C. Cecchi-Pestellini<sup>2</sup>, C.-H. Huang (黃朝暉)<sup>1</sup>,  
Y.-J. Chen (陳俞融)<sup>1</sup>, and G. M. Muñoz Caro<sup>3</sup>

<sup>1</sup>Department of Physics, National Central University, Jhongli City, Taoyuan County 32054, Taiwan

<sup>2</sup>INAF—Osservatorio Astronomico di Palermo, P.za Parlamento 1, I-90134 Palermo, Italy

<sup>3</sup>Centro de Astrobiología (INTA-CSIC), Carretera de Ajalvir, km 4, Torrejón de Ardoz, E-28850 Madrid, Spain  
[angelnnrita@gmail.com](mailto:angelnnrita@gmail.com)

### Abstract

The H<sub>2</sub>O:CO:NH<sub>3</sub> ice mixtures were irradiated by monochromatic (550 and 900 eV) and broadband X-rays (250–1250 eV) to investigate the photodesorption with different photon fluxes and energies. The photodesorbing signals are detected by a quadrupole mass spectrometer (QMS), with dominant species CO ( $m/z = 28$ ), H<sub>2</sub> ( $m/z = 2$ ), and H<sub>2</sub>O ( $m/z = 18$ ). The desorption signals with higher charge to mass ratios for example  $m/z = 69$  (C<sub>3</sub>H<sub>3</sub>NO or C<sub>4</sub>H<sub>5</sub>O or C<sub>4</sub>H<sub>7</sub>N), come from photon-products, and these complex gaseous molecules have been observed in protoplanetary disks. During X-ray bombardment, the desorbing species correspond to the photo-products, which were observed by Fourier transform infrared (FTIR). In this study, we found the photodesorption yields of CO are related to the intensity of OH dangling bond which represents the surface area of pores, and decreasing as the number of absorbed photons increases. The trend of photodesorption yield of CO ice as a function of absorbed photons can be fitted by an equation composed of three exponential terms, which is related to the number of weakly bound species on the surface.

**Keywords** - *ISM: molecules – X-rays: ISM – astrochemistry*